Design Optimization Considerations for MROI

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On behalf of MROI team in NM, USA and Cambridge, UK

Overview of Talk

- Quick Introduction to MROI
- Design Philosophy
- Error Budget
- Subsystem Examples
- Recent Updates
- Quick Movie
- Conclusions



Magdalena Ridge Observatory Interferometer

Federally funded 2000-2011
EIS completed in 2003
Two facilities at MRO
Fast-tracking 2.4m
NIR/Optical 10-element interferometer

MRODis 10x 1.4m movable afoca telescopes in equilateral \ configuration **Optical and near-IR** operation Baselines from 7.8 to 343m **Design optimized for** imaging faint/complex targets



Interferometric Considerations

- Sensitivity
- Speed of Data Collection
 Expert/Novice User Base
- Accuracy
- Precision
- Scalability
- Mobility

- Polarization Purity
- Astrometry
- Imaging
- High Resolution Spectra
- Nulling \bullet
- Costs



MROI Considerations

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MROI Innovations

- Sensitivity = many components in overall design – factor of a few in most cases
- Speed of data collection = simple architecture with very few moving components
- Scalability/mobility = site location & design of building/infrastructure
- Polarization purity = telescopes & beam train
- Imaging = number of telescopes, FT architecture, separated FT/science



Conceptually

Three subsystems – very few mirrors to adjust





Start with an Error Budget

Assign *achievable* errors (factors of a few improvement) to each component of the overall system, and then "hold the line" on



OHP Workshop

Array Layout

- Equilateral Y Configuration
 - 5:1 within a configuration
 - 44:1 overall angular scaling
- Multiple Telescopes
 UV points & CP triangles
- Speed data collection
- Scalability
- Imaging
- Cost infrastructure



aseline coordinate

Telescope coordinate

۲ (m)



Telescopes

- Alt-alt design
 - 3 mirrors
 - 1 reflection > 30°
 - OPD vibrations <40nm
 RMS *tested*
 - <350um radial pupil motion *tested*
- Sensitivity, Scalability, Mobility, Polarization Purity, Imaging





Pirnay et al. 2008 SPIE

Agdalena Ridge Observatory

FTT and Beam Relay Systems

- Transmissive design
- Located on Nasmyth platform – commands 2^{ndary}
- EM CCD
- TT error of 60.8mas RMS at m_v = 16 in 0.7" seeing
- Sensitivity, Speed of Data Collection
- Cost





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- Vacuum transport large beams (cheap)
- No longitudinal dispersion correction needed
- Multiple wavelengths simultaneously
- Sensitivity
- Cost



Delay Lines

- Single stroke system with novel features
 - 380m continuous delay compensation
 - Vacuum vessel and track are same component
 - No "on-board" cabling
 - COTS laser metrology
 - $-\lambda/40$ RMS DL jitter over 2t_o *tested* on 25m



Fischer et al. 2010 SPIE

13

Haniff et al. 2008 SPIE



Delay Lines

- Sensitivity, Speed of Data Collection, Polarization Purity, Imaging
- Cost (\$275K after NRE)



FT - ICONN

- Nearest-Neighbors for all 10 beams
 - Maximizes per baseline SNR
 - Accepts 5 beams per dewar
- No adjustments, precision plates
- Reflective outside, refl/trans inside
- Stays aligned for days in campus lab
- Sensitivity, Scalability, Speed of Data Collection
- Cost







Recent MROI Updates

- UTM1 shipped last week
- DL1 cart delivered few months back and lab tested; will be installed on Ridge in 2 months
- Visitor/maintenance building on Ridge under construction
- First BR system installed outside building
- University searching for more \$\$ Fed & State



Last full talk at Flagstaff meeting in Feb 2013 if you want more MROI details.

MOVIE HERE!!

Conclusions

- Sensitivity = 14th mag at H for FT
- Polarization purity = 1 reflection >30° AOI
- Scalability = Entire design for 10 telescope array add on components as \$ comes
- Mobility = 28 pads, 4 "intrinsic" configurations
- Speed of imaging = 10 min/calibrated UV snapshot
 Snapshot Imaging Machine!

See Ifan Payne if you are interested in joining.



Thank you for your attention!

